

Observations of long-period rotational ground motions: from ambient noise to Earth's free oscillations

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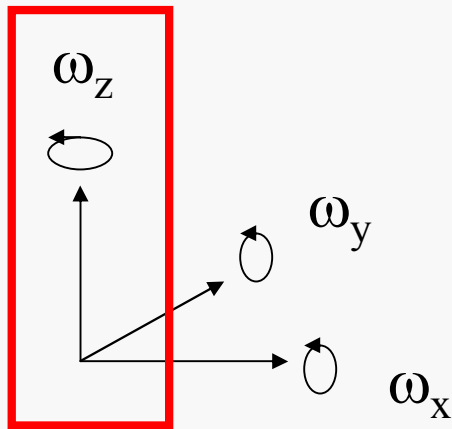
⁵ now at Utrecht University, Netherlands

⁶ IGPP, UCSD, USA

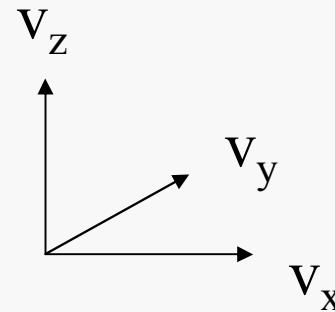




$$\begin{pmatrix} \omega_x \\ \omega_y \\ \omega_z \end{pmatrix} = \frac{1}{2} \nabla \times \underline{\mathbf{v}} = \frac{1}{2} \begin{pmatrix} \partial_y v_z - \partial_z v_y \\ \partial_z v_x - \partial_x v_z \\ \partial_x v_y - \partial_y v_x \end{pmatrix}$$



Rotation rate
Rotation sensor

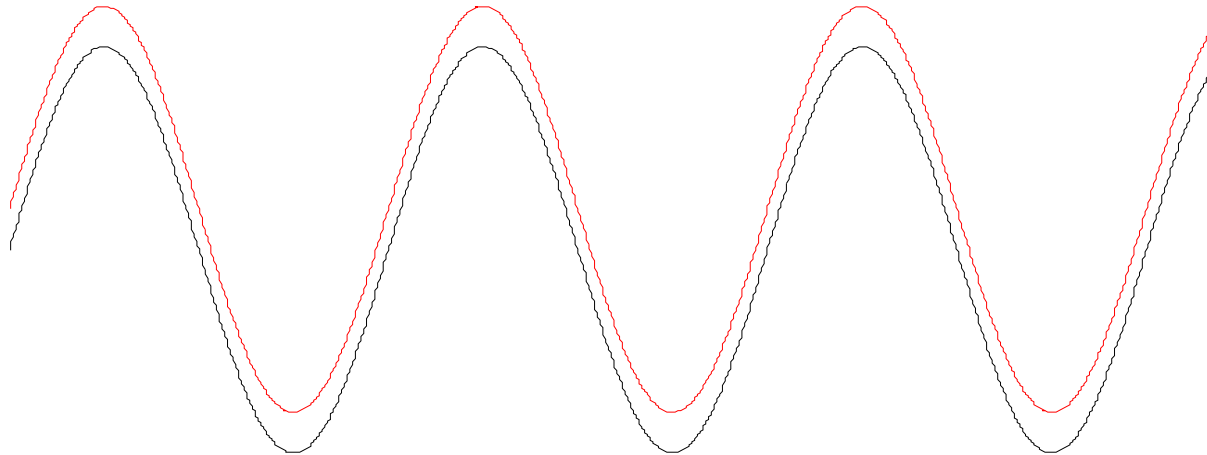


Ground velocity
Seismometer



Plane transversely polarized wave propagating in x-direction with phase velocity c

rotation rate – **transverse acceleration**

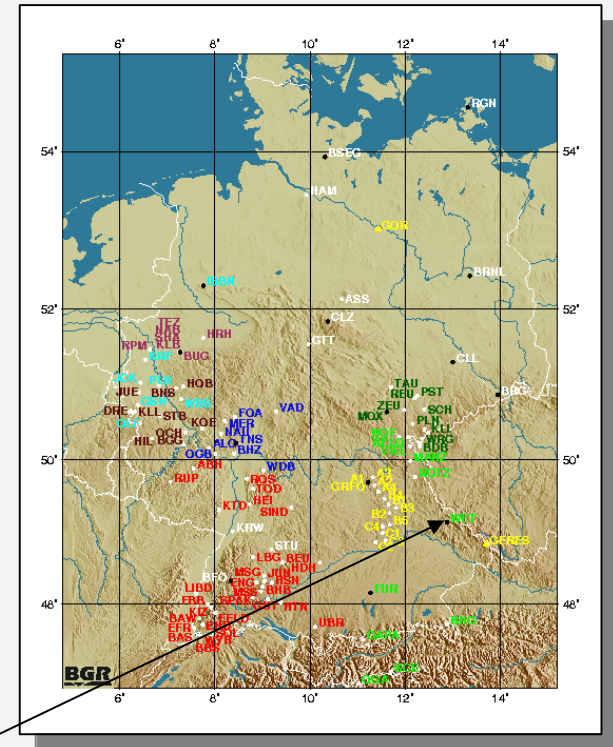
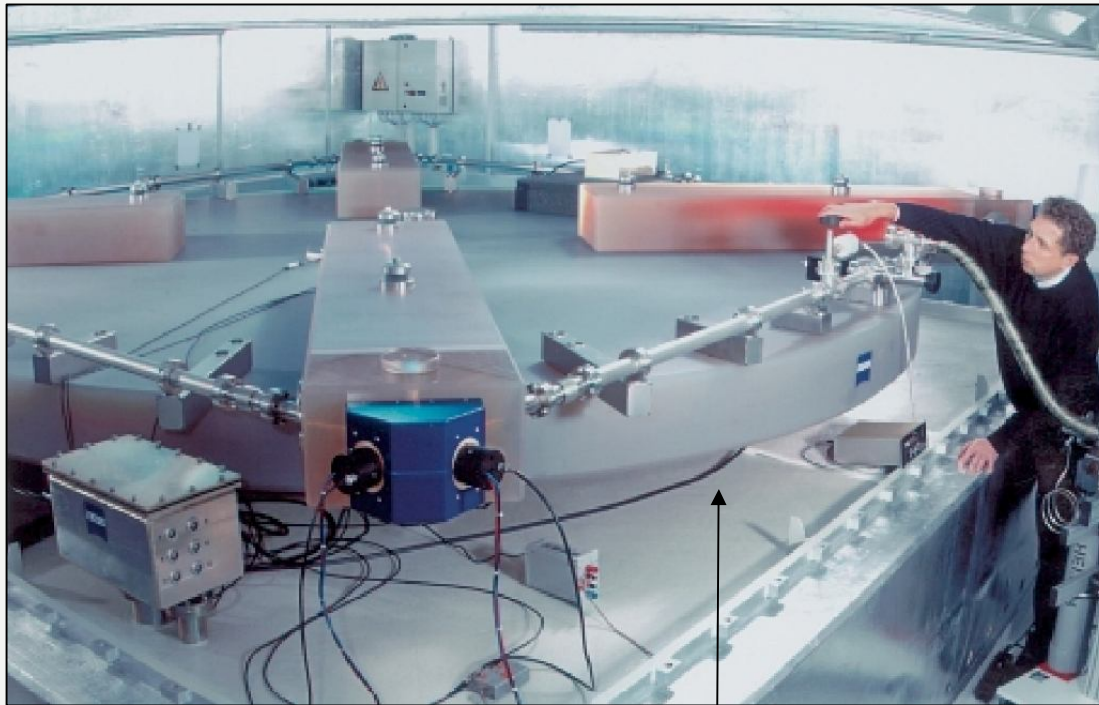


$$a(x, t) / \dot{\Omega}(x, t) = -2c$$

Rotation rate and acceleration should be **in phase** and the **amplitudes scaled by two times the horizontal phase velocity**



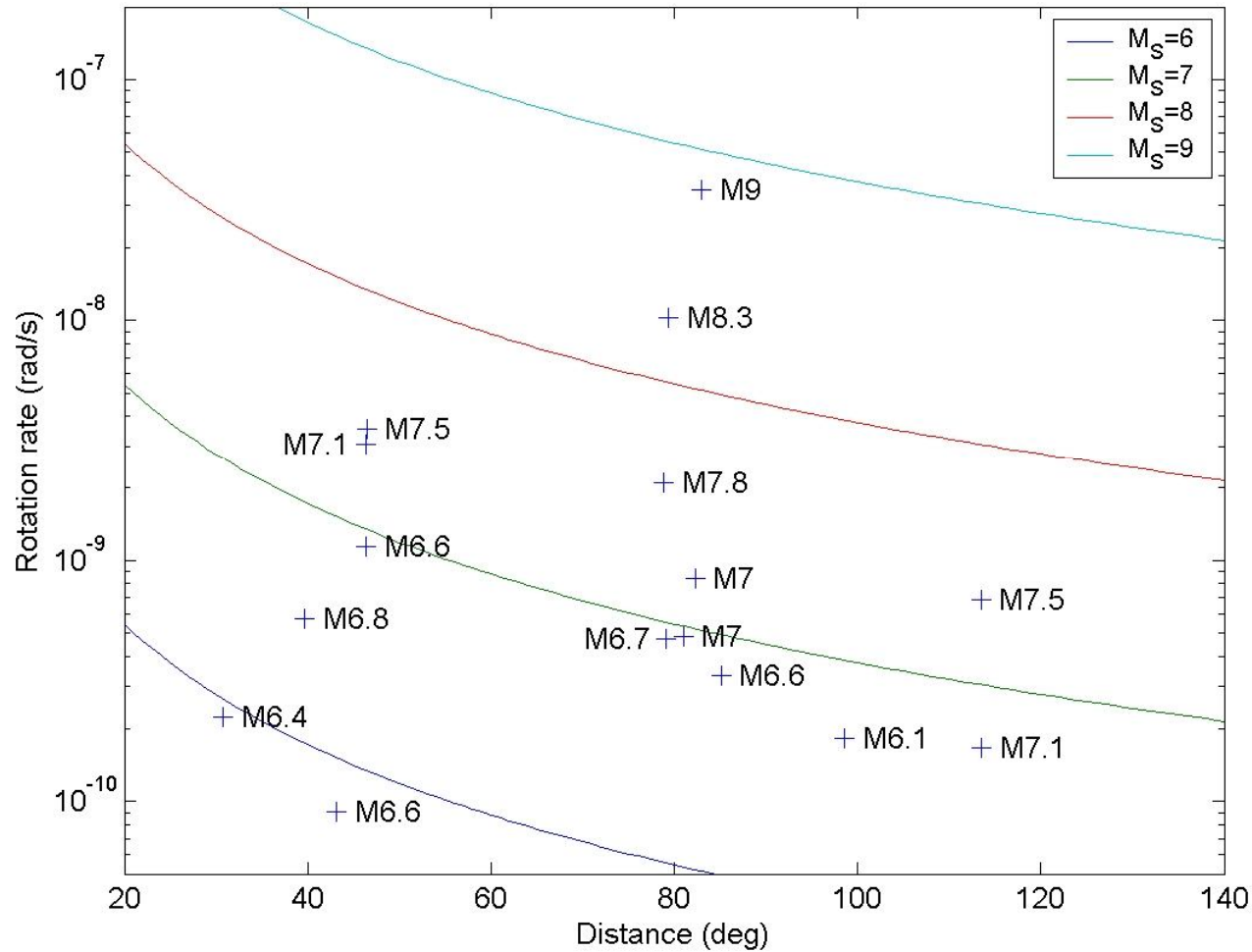
Instrument principle described in Schreiber et al., BSSA, 2009, special issue.



Ring Laser

Rotation rate amplitudes: what to expect?

Observations



M=9

M=8

M=7



Plane wave analysis

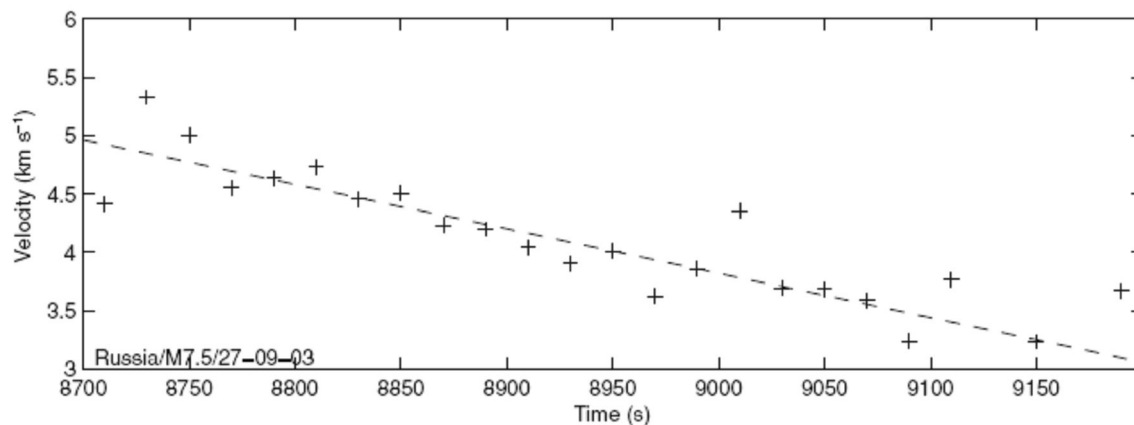
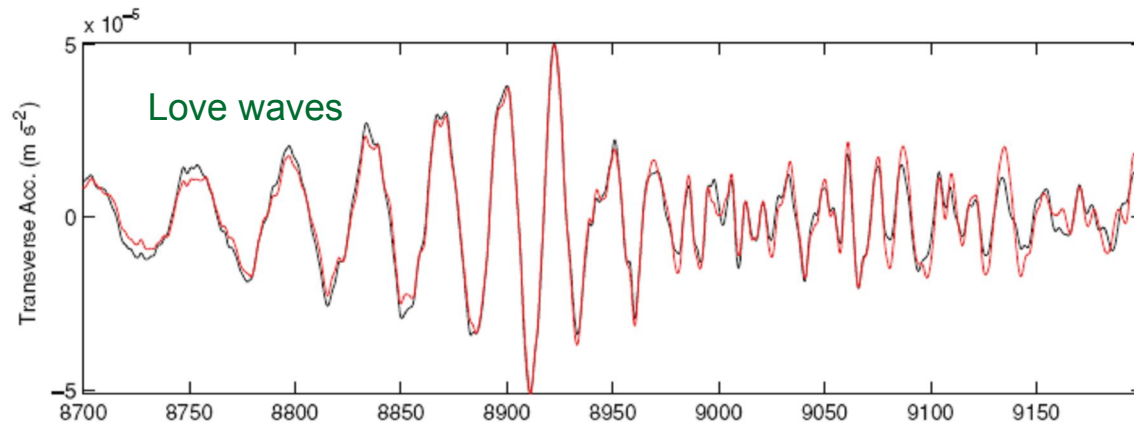
tells us:

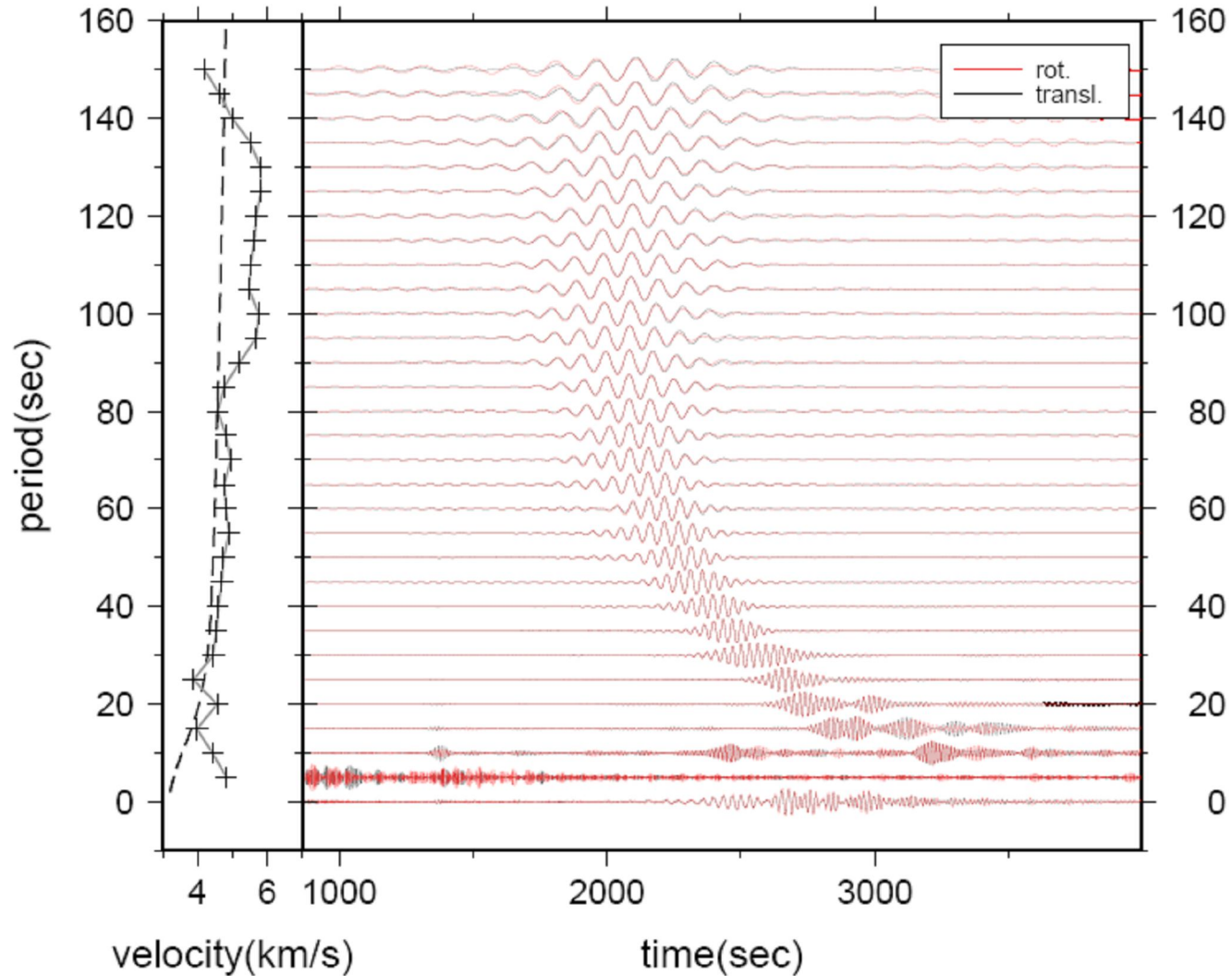
- Waveforms (transverse acceleration and rotation rate) should be in phase
- The ratio between acceleration and rotation rate is proportional to phase velocity
- This can be used for **seismic tomography without travel times**

References (e.g.):

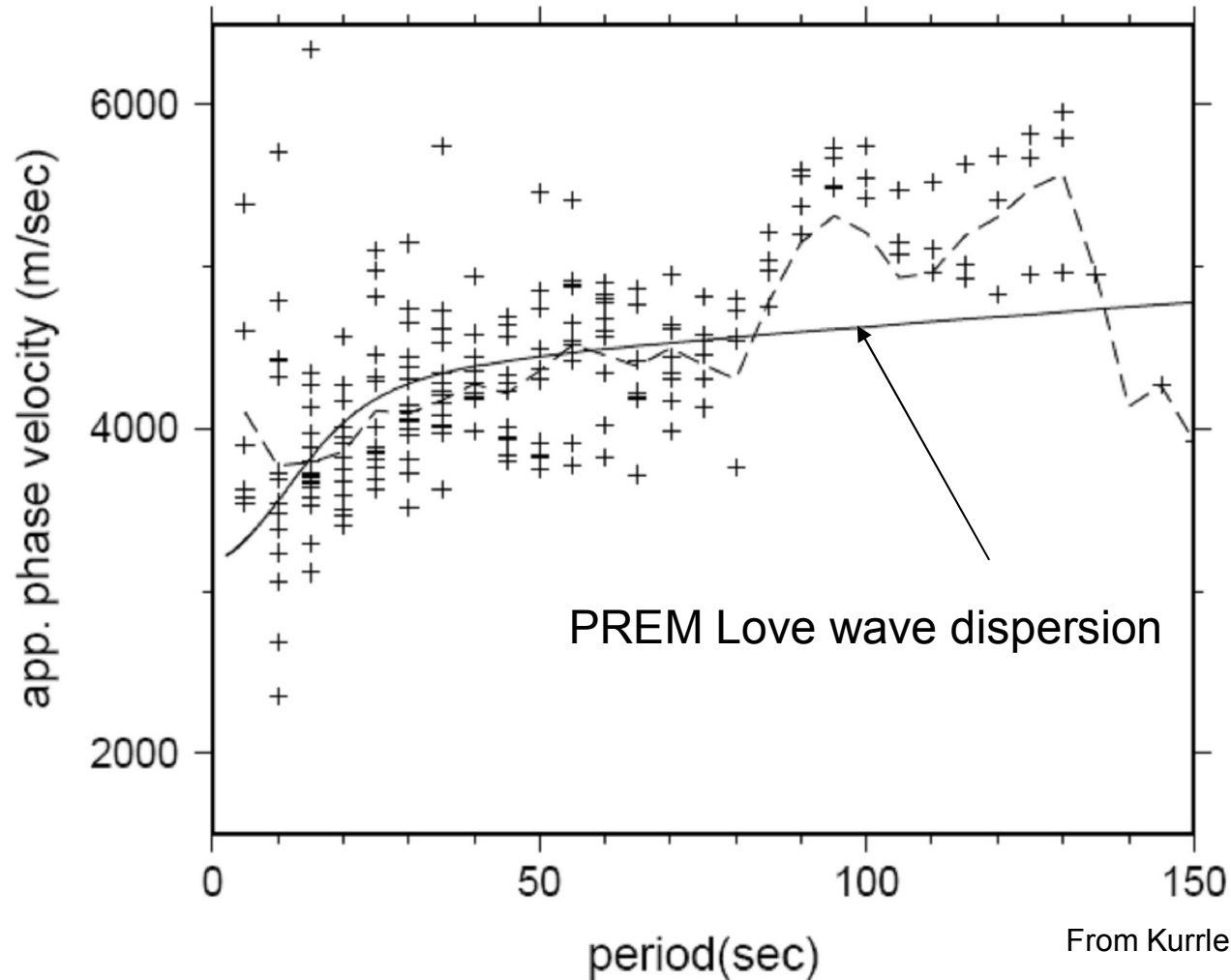
- Pancha et al., GRL, 2000
 Igel et al., GJI 2007
 Ferreira and Igel, BSSA, 2009
 Kurrle et al., GRL, 2010
 Fichtner et al., BSSA, 2009
 Bernauer et al., Geophysics, 2009

rotation rate – transverse acceleration



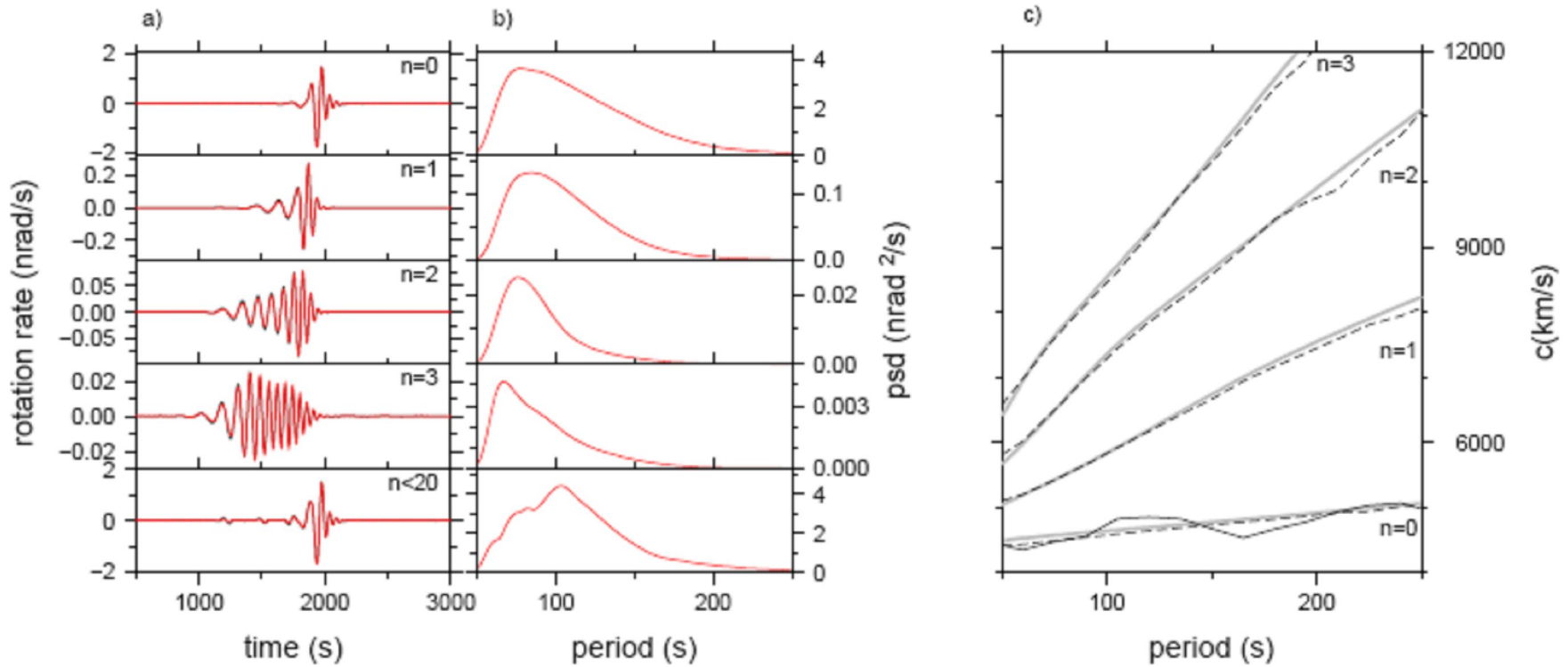


Mw = 7.4 Kuriles earthquake on Jan 15, 2009



From Kurrle et al., GRL, 2010

Why is it so difficult?



... higher mode Love wave contributions can spoil it!

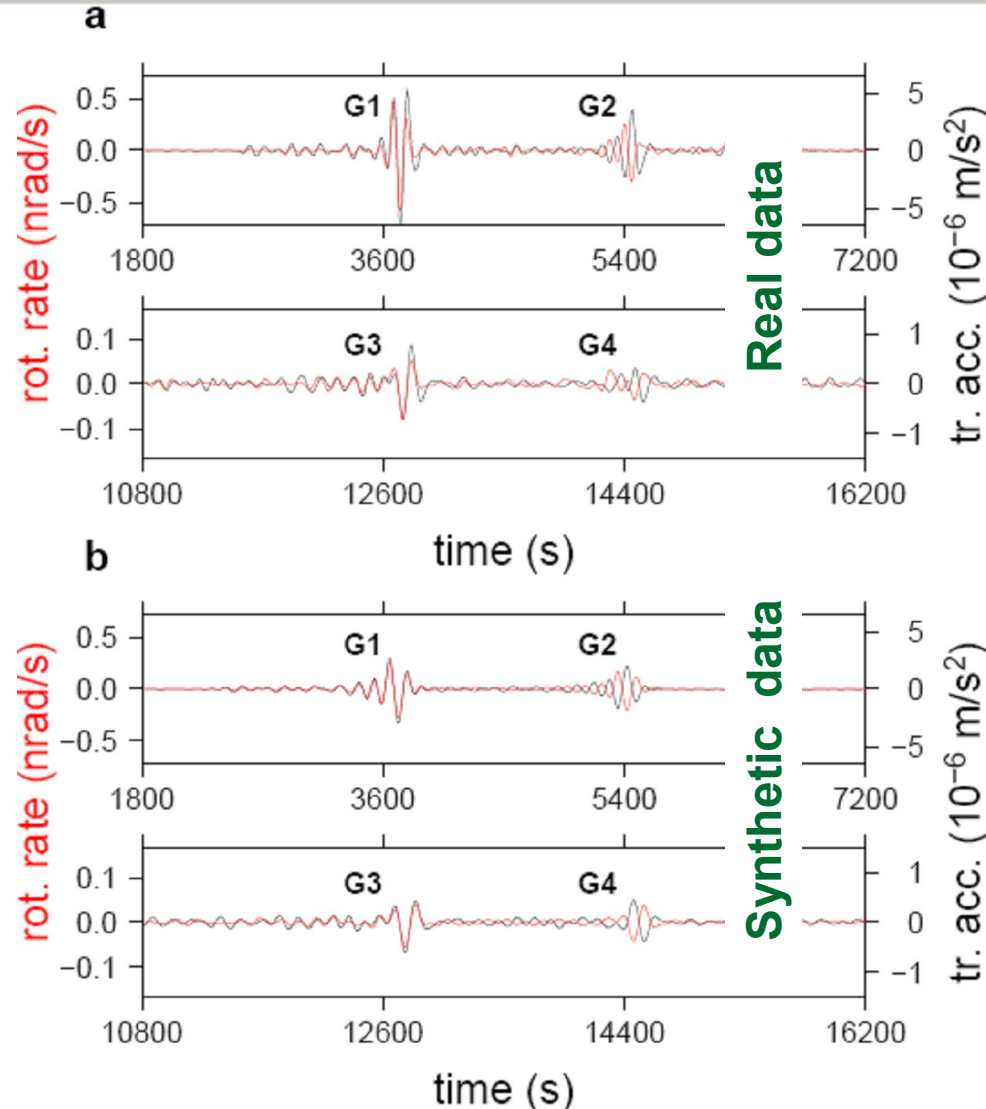


Because at low frequencies ($f < 5$ mHz) horizontal seismometers are limited by tilt noise, there exists the possibility for obtaining superior torsional mode spectra with ring lasers provided that their self noise is further reduced.

Widmer et al., BSSA, 2009



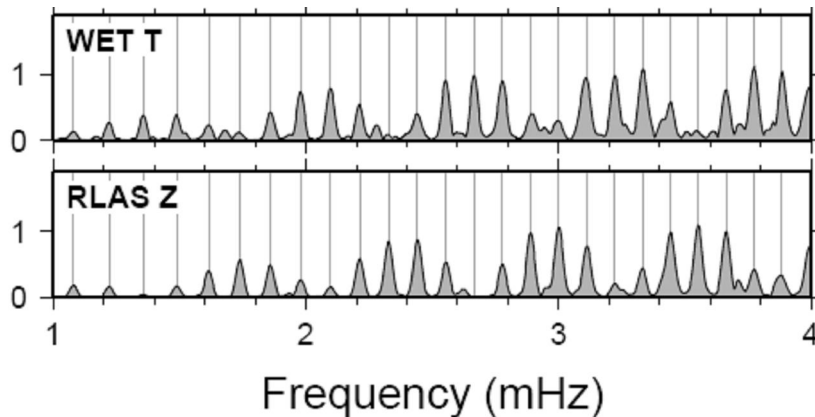
Magnitude 8.1 - SAMOA ISLANDS REGION
2009 September 29 17:48:10 UTC



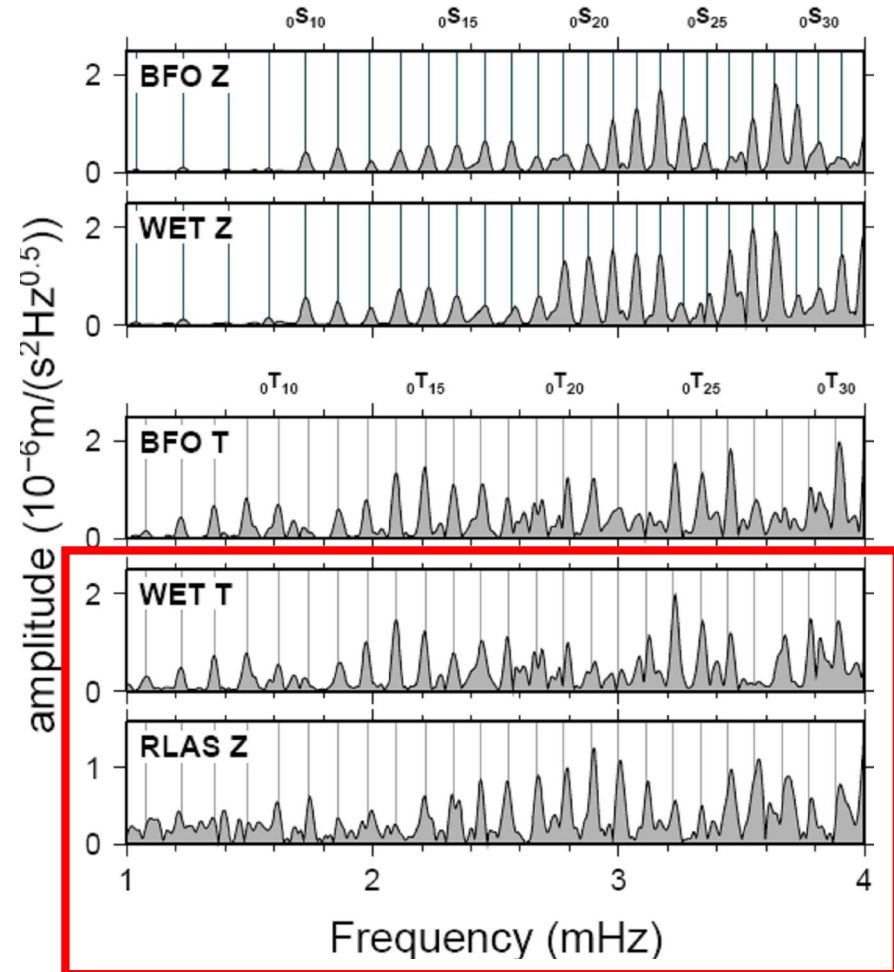


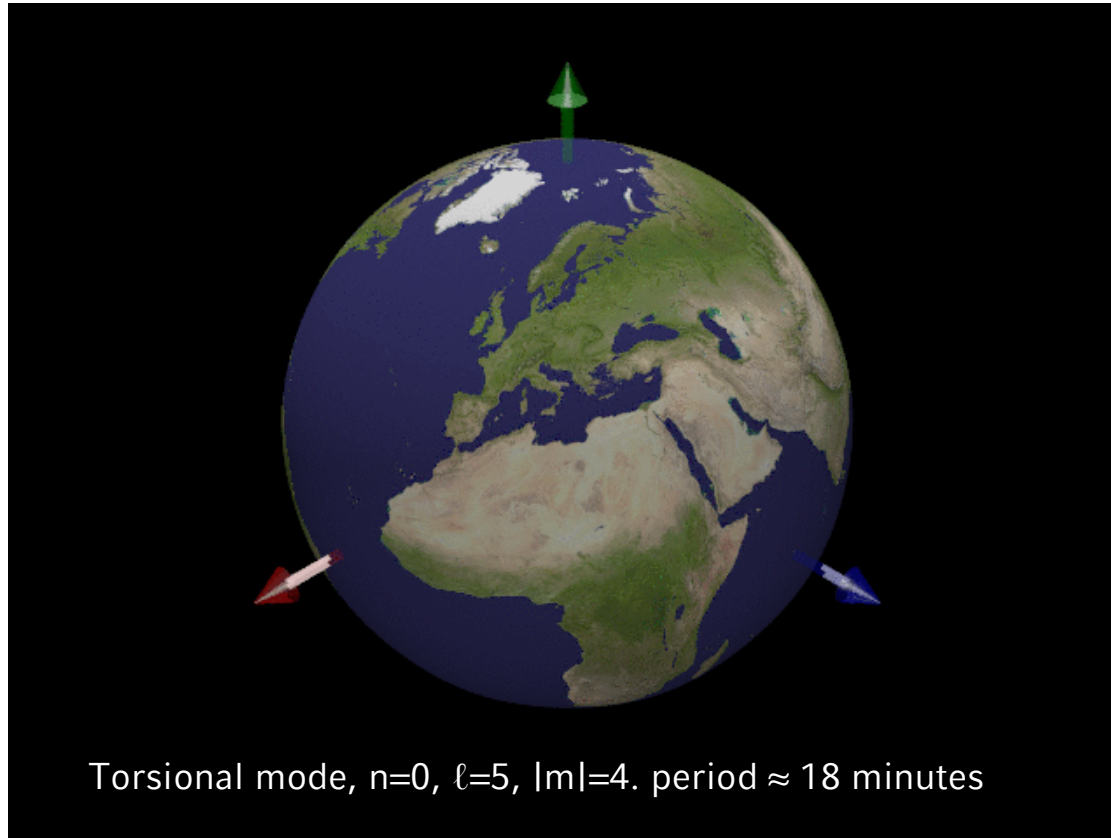
- 16 hr time window
- Comparison with BFO
- Hanning taper

Synthetic spectra (normal modes)



Observations

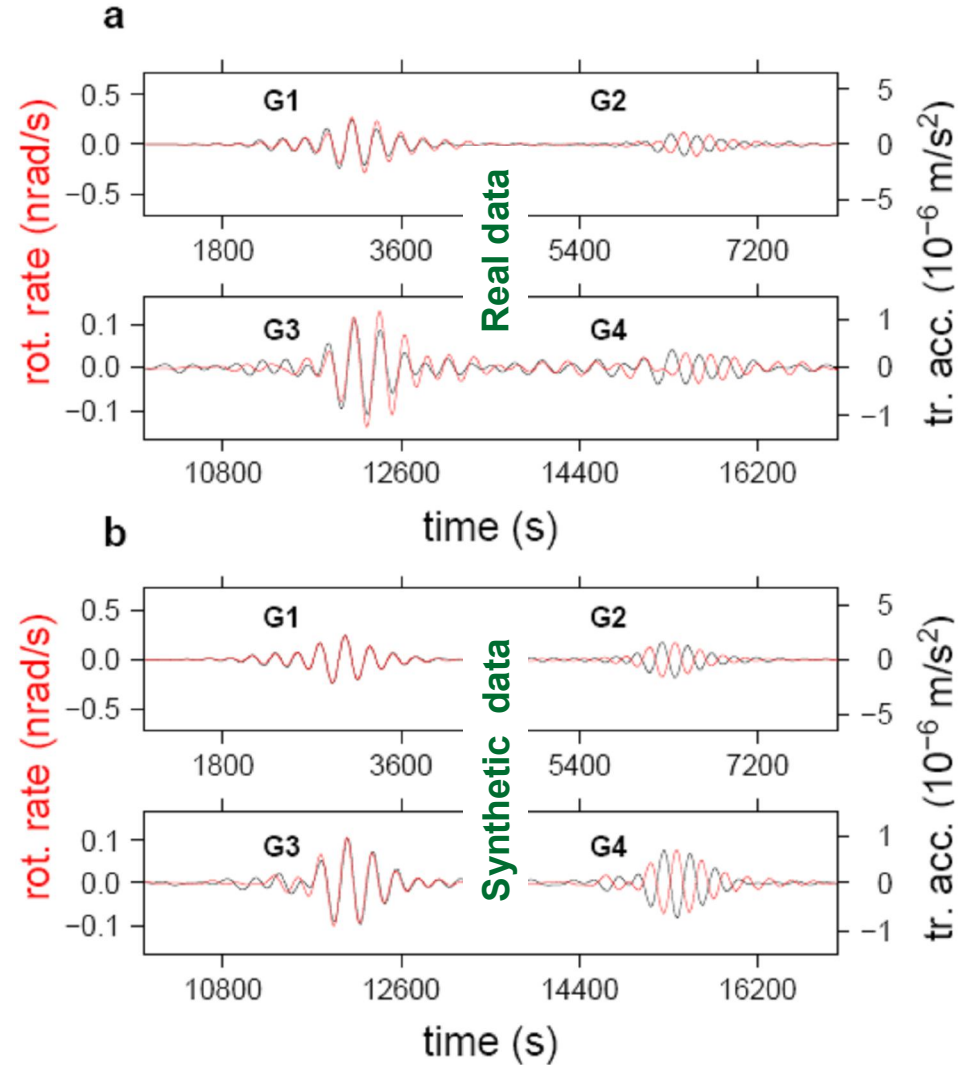




Source: <http://icb.u-bourgogne.fr/nano/MANAPI/saviot/terre/index.en.html>



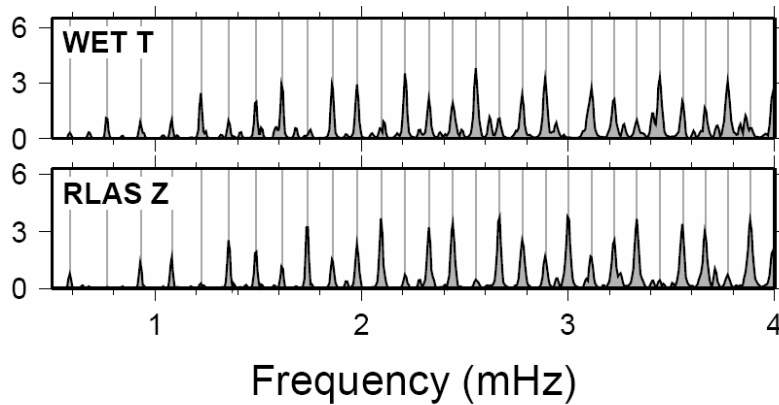
Magnitude 8.8 – OFFSHORE MAULE, CHILE
2010 February 27 06:34:14 UTC



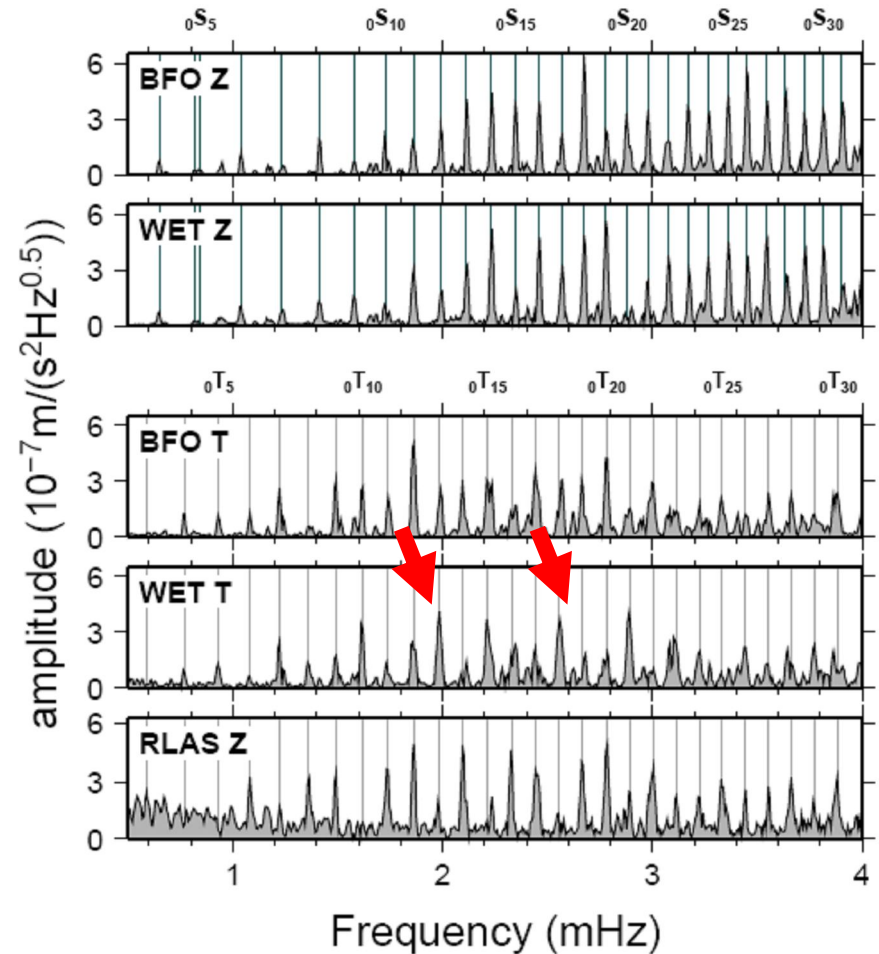


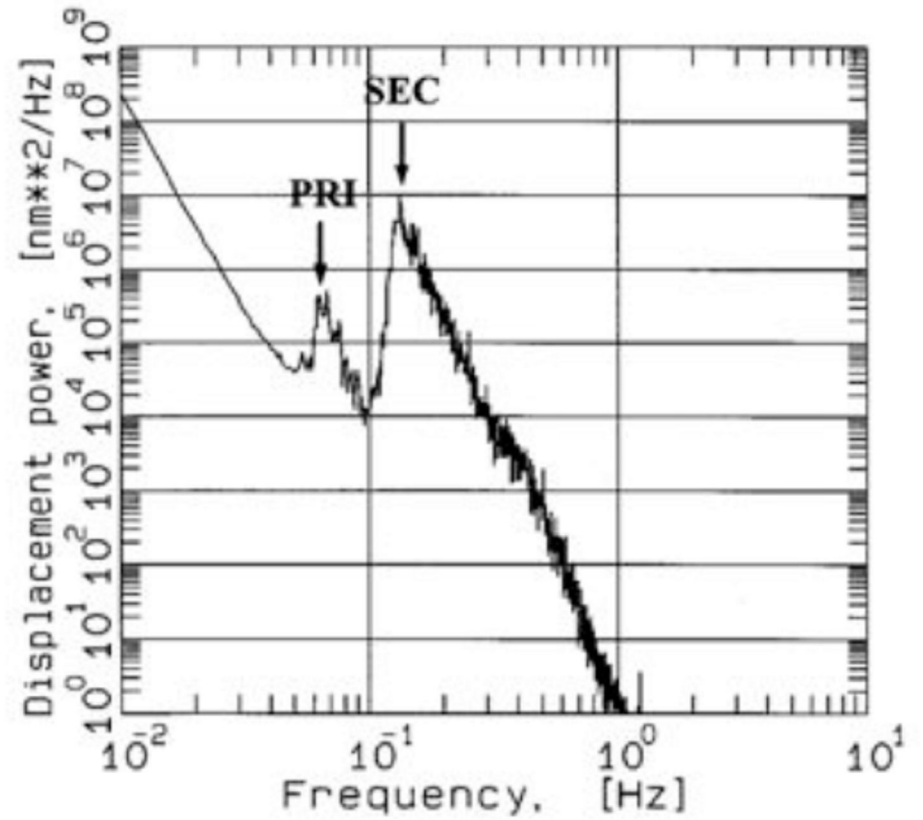
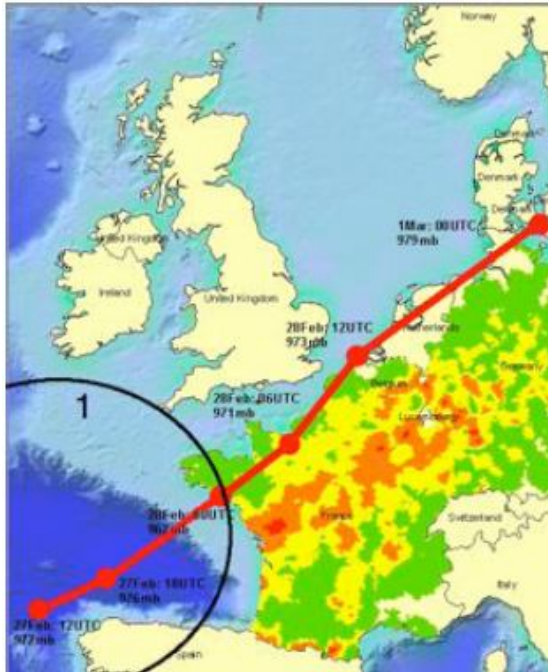
- 36 hr time window
- Comparison with BFO

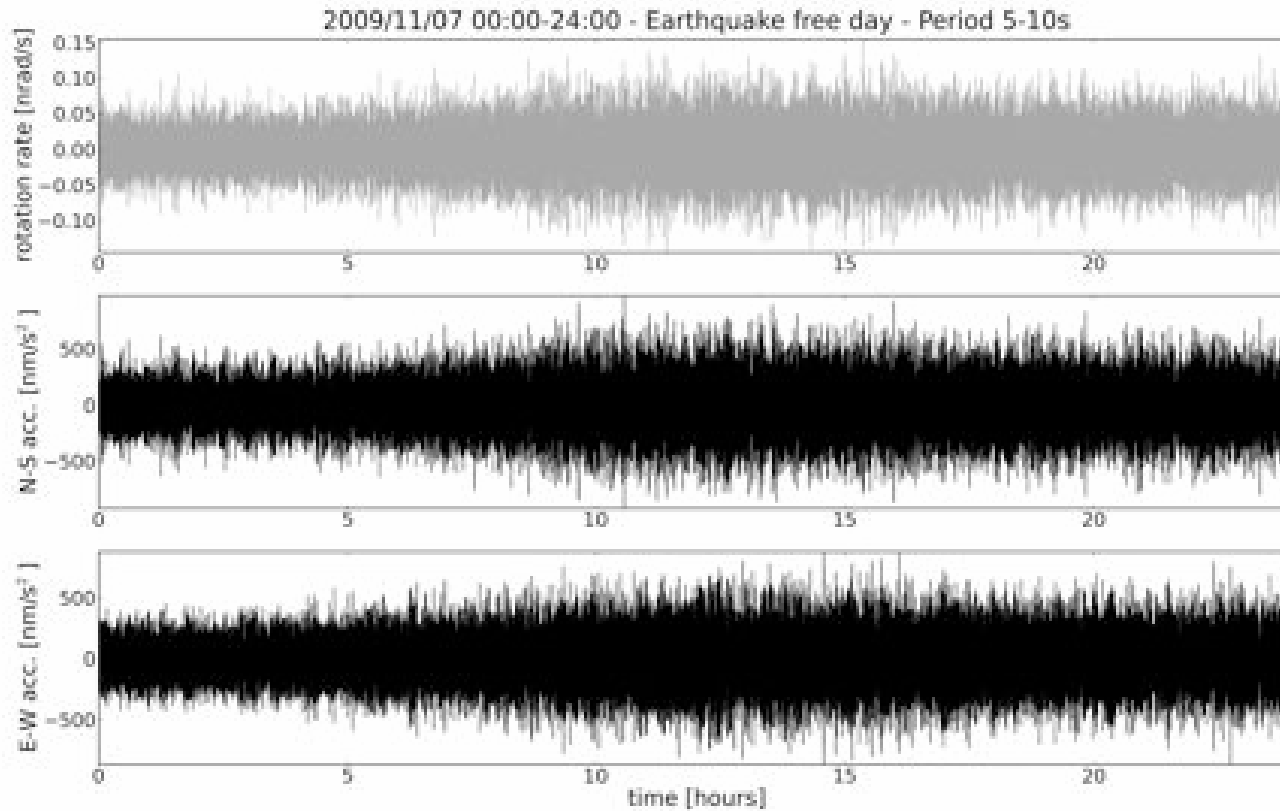
Synthetic spectra (normal modes)

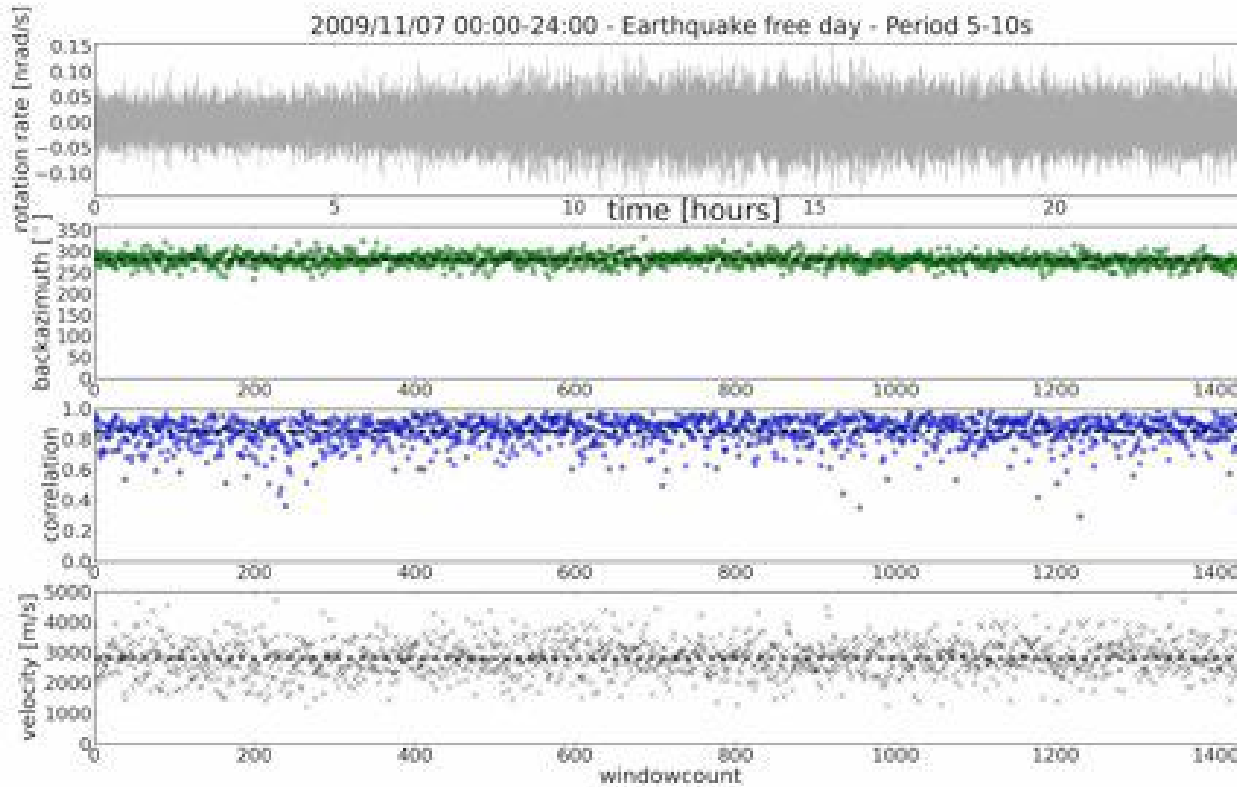


Observations









Rotation rate

Back azimuth

Correlation coefficient

Phase velocity

Total length 24 hours, window length 1 minute



- Observations of **Earth's free oscillations** (torsional modes) are complementary (to standard observations), we would like to go 3C with rotational measurements
- The detection of Love waves in ambient noise might open new possibilities for **correlation type techniques**
- **Pure broadband tilt** (uncontaminated by transverse accelerations) **has never been observed!** (more precise seismometer recordings, gravitation wave detection, displacement histories, health of buildings, etc.)